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'Ghost Particles' Erupting from Antarctica Could Shatter Standard Physics Rare Cosmic Ray Caught in a Cloud Chamber **CITA 79: Supernova remnants and the effect of efficient cosmic ray acceleration LA FÊTE DE LA SCIENCE À l'IAP** *Strange Unexplained Cosmic Rays (ft PhD Comics) What are Cosmic Rays?*

Cosmic Ray Interactions Propagation And

Cosmic-ray propagation and interactions in the Galaxy Andrew W. Strong¹, Igor V. Moskalenko², Vladimir S. Ptuskin³ 1. Max-Planck-Institut für extraterrestrische Physik, Postfach 1312, 85741 Garching, Germany. aws@mpe.mpg.de 2. Hansen Experimental Physics Laboratory (HEPL) and KIPAC, Stanford University, Stanford, CA 94305, U.S.A. imos ...

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We survey the theory and experimental tests for the propagation of cosmic rays in the Galaxy up to energies of 10¹⁵ eV. A guide to the previous reviews and essential literature is given, followed by an exposition of basic principles. The basic ideas of cosmic-ray propagation are described, and the physical origin of its processes are explained. The various techniques for computing the ...

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Cosmic Ray Propagation in Space Plasmas. Pages 109-403. Nonlinear Cosmic Ray Effects in Space Plasmas. Pages 405-494. ... Cross section Galaxy Neutrino acceleration astrophysics cosmic rays in space plasma interactions nonlinear effects propagation thermodynamics . Authors and affiliations.

Cosmic Ray Interactions, Propagation, and Acceleration in

... Cosmic ray propagation and interactions in the Galaxy V.N.Zirakashvili Puskov Institute of Terrestrial Magnetism, Ionosphere and Radiowave Propagation, Russian Academy of Sciences, 142190 Troitsk Moscow, Russia Abstract Cosmic ray propagation in the Galaxy is shortly reviewed. In particular we consider the self-consistent models of CR propagation.

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Astrophysics of Galactic Cosmic Rays (15) gives a valuable overview of the experimental data and theoretical ideas as of 2001. The biannual International Cosmic Ray Conference proceedings (many of which can be found at <http://adswww.harvard.edu/>) are

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Scientists can reconstruct a historical record of cosmic-ray flux by measuring the concentration of the isotope beryllium-10 in rocks. The beryllium-10 is a product of cosmic-ray interactions in the atmosphere, but the timescale of its radioactive decay means that such records only go back a few million years.

Physics - Cosmic-Ray Time Capsules

The resulting general transport theory relates the motion of cosmic rays, statistically, to the solar-wind velocity and magnetic field. Application of the theory both to the modulation of galactic cosmic rays by the solar wind and to the propagation of solar cosmic rays is discussed in detail.

Propagation of cosmic rays in the solar wind - Jokipii ...

Cosmic rays ionize the nitrogen and oxygen molecules in the atmosphere, which leads to a number of chemical reactions. Cosmic rays are also responsible for the continuous production of a number of unstable isotopes in the Earth's atmosphere, such as carbon-14, via the reaction: $n + {}^{14}\text{N} \rightarrow p + {}^{14}\text{C}$

Cosmic ray - Wikipedia

We survey the theory and experimental tests for the propagation of cosmic rays in the Galaxy up to energies of 10¹⁵ eV. A guide to the previous reviews and essential literature is given, followed by an exposition of basic principles. The basic ideas of cosmic-ray propagation are described, and the physical origin of its processes is explained.

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We study the effects of drift motions and the advection by a Galactic wind on the propagation of cosmic rays in the Galaxy. We employ a simplified magnetic field model, based on (and similar to) the Jansson-Farrar model for the Galactic magnetic field. Diffusion is allowed to be anisotropic. The relevant equations are solved numerically, using ...

The effects of drift and winds on the propagation of ...

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We discuss the propagation of Extremely High Energy (EHE) cosmic ray protons and neutrinos. Their extremely high energies involve the interactions with the cosmological relic radiations of photons and neutrinos, which result in the modifications of the energy spectra of EHE particles after their propagation. The GZK cutoff for protons and the dip and steepening for neutrinos.

Propagation of cosmic rays and neutrinos through space

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of the cosmic radiation in the GeV range depends both on the location and time. There are four different ways to describe the spectra of the components of the cosmic radiation: (1) By particles per unit rigidity. Propagation (and probably also acceleration) through cosmic magnetic fields depends on gyroradius or magnetic rigidity, R , which is

30.CosmicRays

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